

On the Depolarization of the Observed Emission From Radio Relics

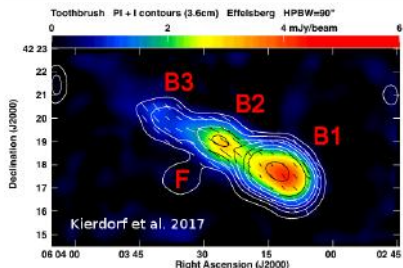
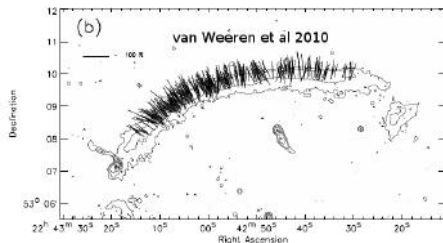
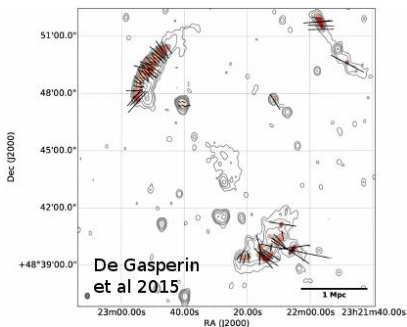
Denis Wittor

with F. Vazza, Matthias Hoeft and M. Brüggén

21.03.2018



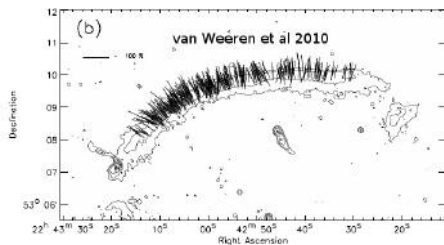
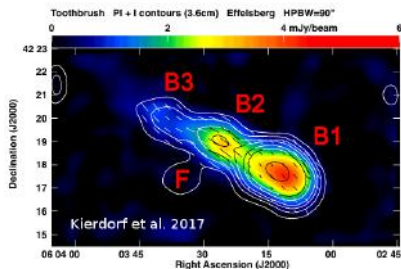
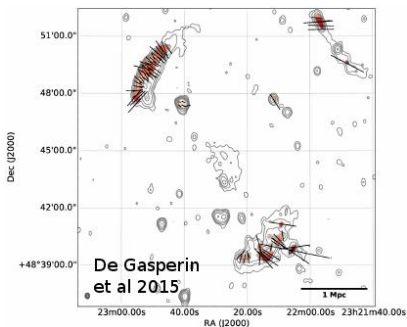
POLARIZATION OF RADIO RELICS



Polarization of Radio Relics

- highly polarized 10 – 50 %
- a lot of observations (Bonafede et al. 2009, van Weeren et al. 2010, Kale et al. 2012, de Gasperin et al. 2015, Kierdorf et al. 2017, ...)
- but no simulations (Skillman et al. 2013)

POLARIZATION OF RADIO RELICS



Polarization of Radio Relics

Use polarized emission from radio relics to study the magnetic field properties using numerical simulations.

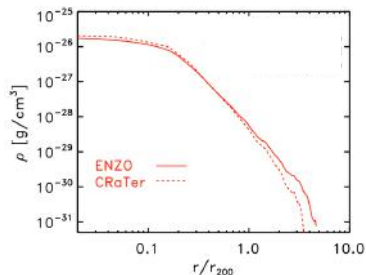
SIMULATIONS

Cosmological Simulations with ENZO

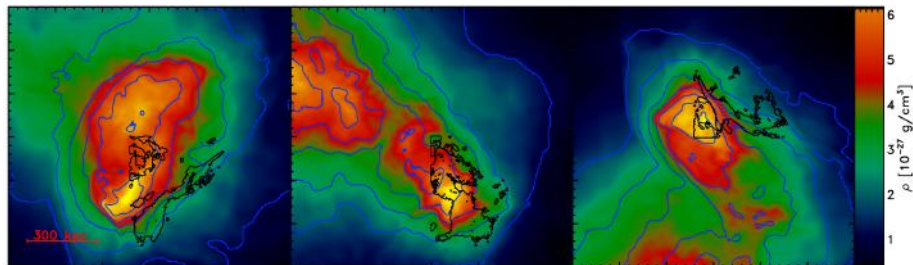
- by Bryan et. al 1997 ... 2014
- dark matter: N-body particle mesh-solver
- baryonic matter: adaptive mesh method

Lagrangian Analysis with CRaTer

- follow baryonic mass passively
- detect shocks using temperature jumps
- compute shock injected cosmic-ray electron spectrum
- e.g. Wittor et al. 2016, 2017



HIGH RESOLUTION RADIO RELIC



ENZO simulation

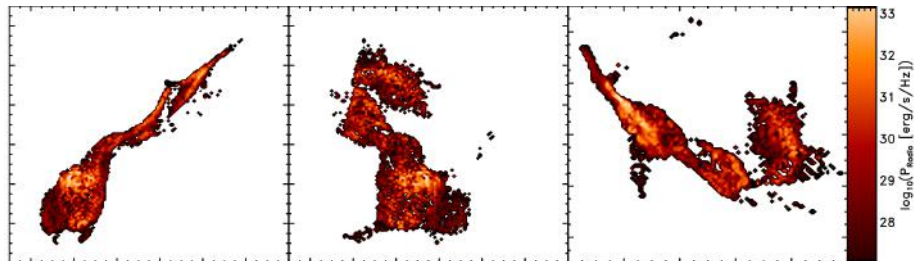
- resolution
 $dx \approx 4 \text{ kpc}$
- timestep
 $dt \approx 4 \text{ Myr}$

CRaTer simulation

- $5 \cdot 10^7$ tracers in a $(1600 \text{ kpc})^3$ volume
- $m = 2 \cdot 10^6 M_{\odot}$
- $5 \cdot 10^5$ shock injected spectrum following Hoeft & Brüggem 2007

color: density
blue: pressure
black: radio

RADIO RELIC CLOSE-UP

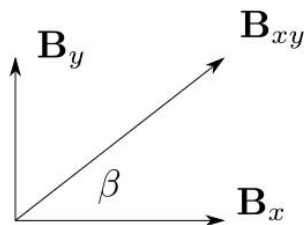
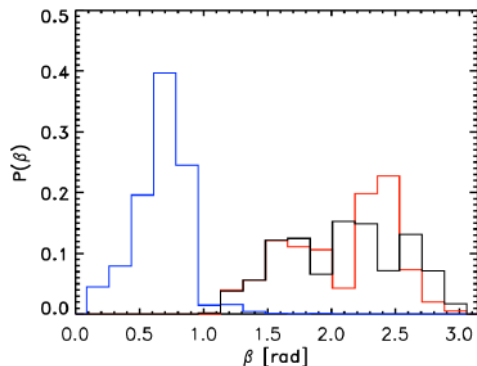


Goal: Polarized Emission (Burn 1966)

$$P_{\text{burn}}(\lambda^2) = \frac{\int_{\text{los}} P_{\text{tot}} \Pi \exp(2i(\beta + RM\lambda^2)) ds}{\int_{\text{los}} P_{\text{tot}} ds}$$

- Π : fractional polarization
- β : intrinsic angle of polarization
- $\Phi\lambda^2$: Faraday Rotation

INTRINSIC ANGLE OF POLARIZATION



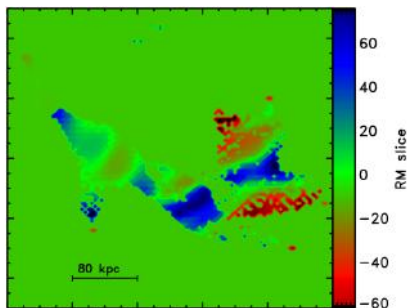
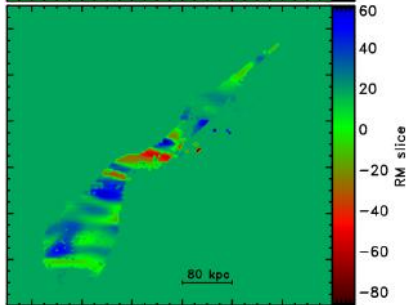
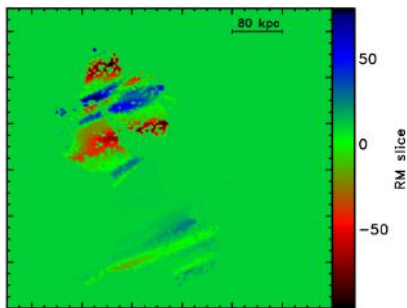
Color: different LoS

black: $0.5 \cdot \pi < \beta < \pi$

blue: $0.5 \cdot \pi < \beta < \pi$

red: $0 < \beta < 0.5 \times \pi$

DEPOLARIZATION



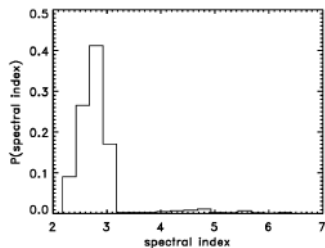
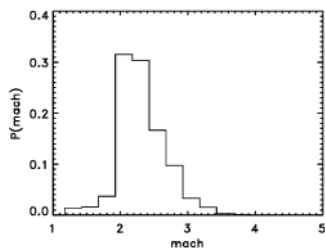
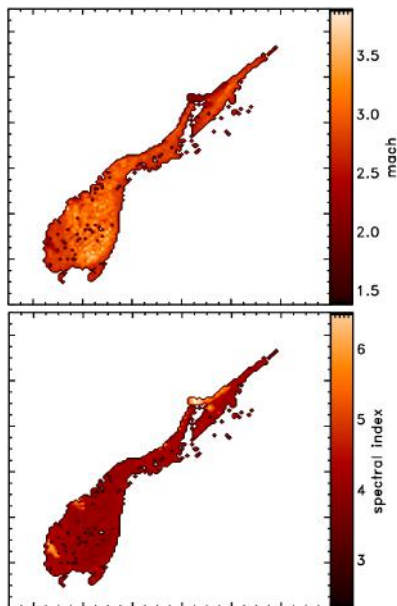
Depolarization

- Faraday Rotation $RM\lambda^2$:

$$RM \propto \int n_e \mathbf{B}_{\parallel} dl \cos$$

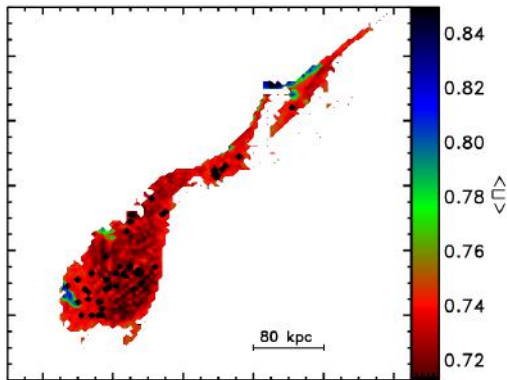
- instrumental depolarization not included

MACH NUMBER AND SPECTRAL INDEX



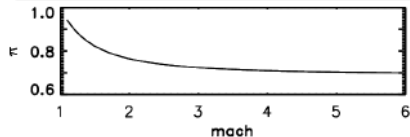
- $\langle M \rangle \approx 2.3$
- $\langle s \rangle \approx 2.8$

FRACTIONAL POLARIZATION

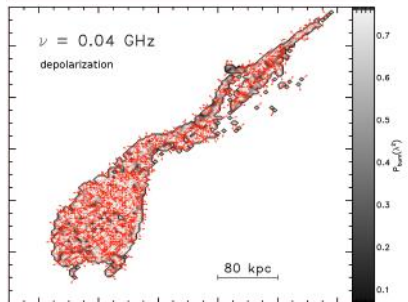
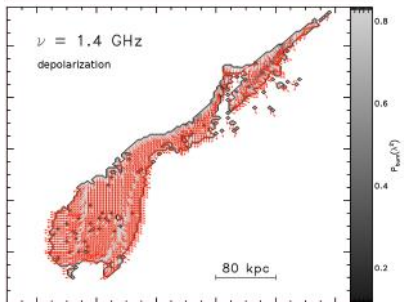
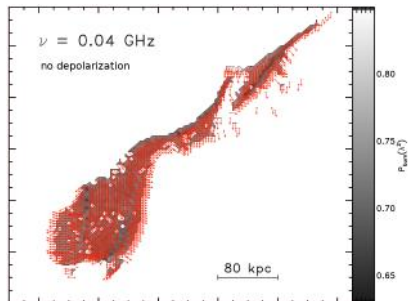
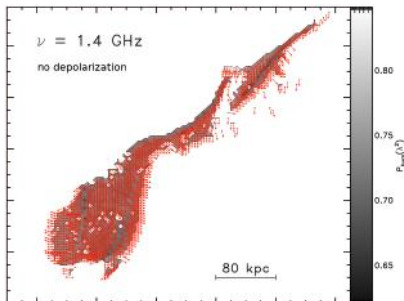


For a distribution of electrons

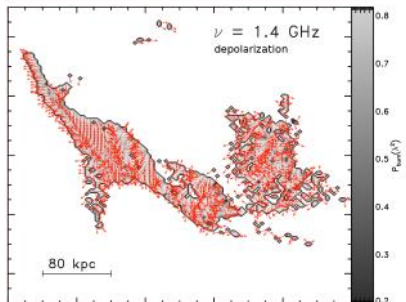
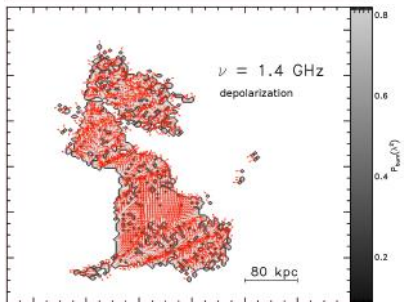
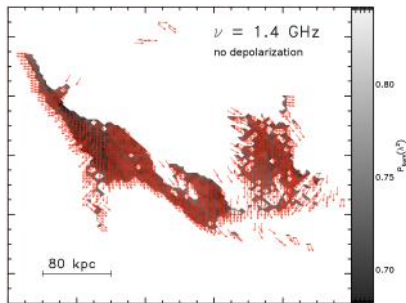
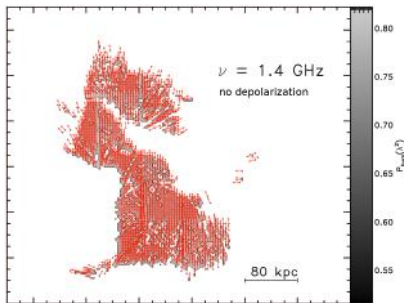
$$\Pi = \frac{3\alpha_{inj} + 3}{3\alpha_{inj} + 7}$$
$$\alpha_{inj} = 2 \frac{M^2 + 1}{M^2 - 1}$$



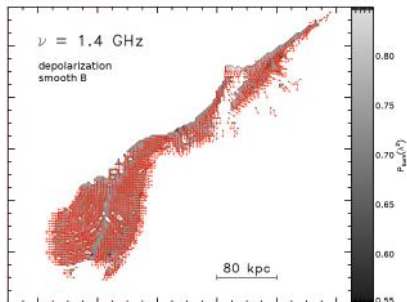
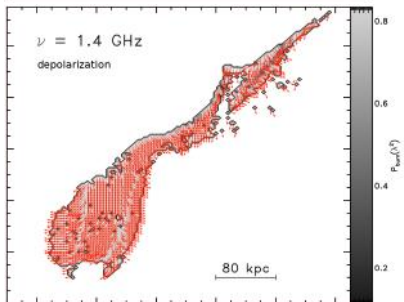
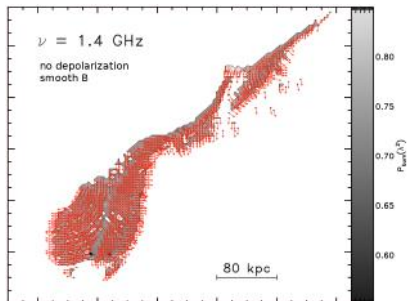
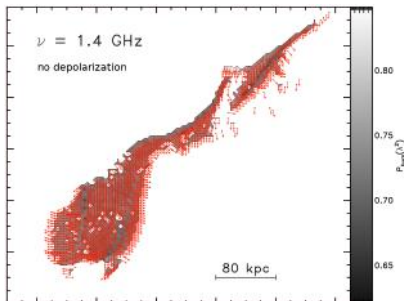
POLARIZED EMISSION



POLARIZED EMISSION: DIFFERENT LOS



POLARIZED EMISSION: SMOOTH B

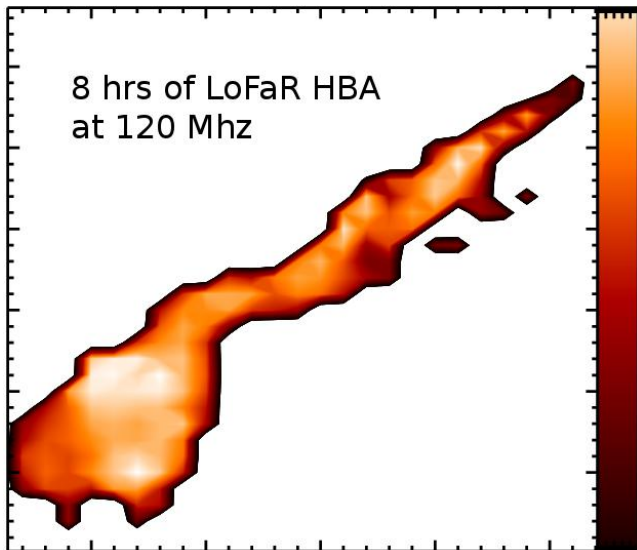


SUMMARY & OUTLOOK

- high resolution simulation of radio relics
 - ▶ high resolution grid simulation with ENZO
 - ▶ analysed using Lagrangian tracer particles
- followed electron spectrum injected by an $\langle M \rangle \approx 2.3$ shock
- computed the polarized emission
 - ▶ independent of the LoS
 - ⇒ due to a laminar magnetic field

- connection to observations
- include cooling of electrons
- increase resolution ($dx \approx 2$ kpc & $dt \approx 2$ Myr)

TEASER



THANK YOU FOR YOU ATTENTION!

ANY QUESTIONS?